R006-23 A 会場:11/7 AM1 (9:00-10:30) 09:15~09:30

あらせ衛星の観測による放射線帯、プラズマ圏の長期変動

#三好 由純 1),栗田 怜 2),片岡 龍峰 3),齊藤 慎司 4),今城 峻 5),堀 智昭 1),中村 紗都子 6),田 采祐 7),寺本 万里子 8),三谷 烈史 9),篠原 育 10),高島 健 11),東尾 奈々 12),笠原 慧 13),横田 勝一郎 14),桂華 邦裕 15),浅村 和史 9),松岡 彩子 16),土屋 史紀 17),熊本 篤志 17),北原 理弘 1),松田 昇也 18),笠原 禎也 19),風間 洋一 20),Wang S.-Y. 21),Tam Sunny W. Y. 22

 $^{(1)}$ 名大 ISEE, $^{(2)}$ 京都大学 生存研, $^{(3)}$ 極地研, $^{(4)}$ 情報通信研究機構, $^{(5)}$ 京大・地磁気センター, $^{(6)}$ IAR&ISEE, $^{(7)}$ 名大 ISEE 研, $^{(8)}$ 九工大, $^{(9)}$ 宇宙研, $^{(10)}$ 宇宙研/宇宙機構, $^{(11)}$ 宇宙研, $^{(12)}$ JAXA, $^{(13)}$ 東京大学, $^{(14)}$ 大阪大, $^{(15)}$ 東大・理, $^{(16)}$ 京都大学, $^{(17)}$ 東北大・理・惑星プラズマ大気, $^{(18)}$ 金沢大学, $^{(19)}$ 金沢大, $^{(20)}$ ASIAA, $^{(21)}$ 台湾・中央研究院, $^{(22)}$ 国立成功大学宇宙・プラズマ科学研究所

Long-term variations of relativistic electrons of the outer radiation belt, plasma sheet electrons and plasmasphere

#Yoshizumi Miyoshi¹⁾, Satoshi Kurita²⁾, Ryuho Kataoka³⁾, Shinji Saito⁴⁾, Shun Imajo⁵⁾, Tomoaki Hori¹⁾, Satoko Nakamura⁶⁾, ChaeWoo Jun⁷⁾, Mariko Teramoto⁸⁾, Takefumi Mitani⁹⁾, Iku Shinohara¹⁰⁾, Takeshi Takashima¹¹⁾, Nana Higashio¹²⁾, Satoshi Kasahara¹³⁾, Shoichiro Yokota¹⁴⁾, Kunihiro Keika¹⁵⁾, Kazushi Asamura⁹⁾, Ayako Matsuoka¹⁶⁾, Fuminori Tsuchiya¹⁷⁾, Atsushi Kumamoto¹⁷⁾, Masahiro Kitahara¹⁾, Shoya Matsuda¹⁸⁾, Yoshiya Kasahara¹⁹⁾, Yoichi Kazama²⁰⁾, S.-Y. Wang²¹⁾, Sunny W. Y. Tam²²⁾

⁽¹ISEE, Nagoya Univ., ⁽²RISH, Kyoto Univ., ⁽³NIPR, ⁽⁴NICT, ⁽⁵WDC for Geomagnetism, Kyoto, Kyoto University, ⁽⁶IAR&ISEE, ⁽⁷ISEE, Nagoya Univ., ⁽⁸Kyutech, ⁽⁹ISAS/JAXA, ⁽¹⁰ISAS/JAXA, ⁽¹¹ISAS, JAXA, ⁽¹²JAXA, ⁽¹³The University of Tokyo, ⁽¹⁴Osaka Univ., ⁽¹⁵University of Tokyo, ⁽¹⁶Kyoto University, ⁽¹⁷Planet. Plasma Atmos. Res. Cent., Tohoku Univ., ⁽¹⁸Kanazawa Univ., ⁽¹⁹Kanazawa Univ., ⁽²⁰ASIAA, ⁽²¹ASIAA, Taiwan, ⁽²²Institute of Space and Plasma Sciences, National Cheng Kung University, Taiwan

The Arase satellite has observed the inner magnetosphere since 2017 and provided long-term observation data about plasma/particles and field/waves, which covers the declining phase of the solar cycle 24 and the following early rising phase of the cycle 25. We investigate spatio-temporal variations in the phase space density that indicate evolution of the acceleration region for energetic electrons. For several magnetic storms, glowing peaks of the phase space density are found during the recovery phase and this fact indicates that the local acceleration processes work for the large flux enhancement of the outer belt electrons. In addition, we find that the location of the glowing peaks in the phase space density varies with the plasmapause location as well as the inner edge of the plasma sheet electrons. This result is consistent with the cross-energy coupling processes driven by wave-particle interactions in which the electron acceleration driven by whistler mode waves is regulated by the ratio of the plasma frequency (fp) and cyclotron frequency (fc) as well as the source electrons for the wave generation.