R006-07

C会場:11/5 PM1 (13:45-15:30)

15:15~15:30

#北村 成寿 1), 天野 孝伸 2), 大村 善治 3), Boardsen Scott^{4,5)}, 北原 理弘 ⁷⁾, 三好 由純 ¹⁾, 中村 紗都子 ⁶⁾, 小 路 真史 1), 加藤 雄人 7), 小嶋 浩嗣 8),Lee Sun-Hee⁹⁾,Gershman Daniel J.⁴⁾, 齋藤 義文 ¹⁰⁾, 平原 聖文 ¹¹⁾, 横田 勝一郎 12),Giles Barbara L.4),Paterson Craig J.¹³⁾,Le Contel William R.⁴⁾, Pollock Olivier¹⁴, Russell James L.¹⁹⁾ Christopher¹⁵⁾,Strangeway Robert J. ¹⁶⁾,Lindqvist Per-Arne¹⁷⁾,Ergun Robert E.¹⁸⁾,Burch $^{(1)}$ 名大・宇地研, $^{(2)}$ 東大・理, $^{(3)}$ 京大・生存圏, $^{(4)}$ NASA/GSFC, $^{(5)}$ メリーランド大学ボルチモア郡校 惑星太陽圏研, $^{(6)}$ 名 大・IAR&ISEE, $^{(7)}$ 東北大・理・地球物理, $^{(8)}$ 京大・生存圏, $^{(9)}$ カトリック大学アメリカ・物理天体, $^{(10)}$ 宇宙研, $^{(11)}$ 名大・ 宇地研,⁽¹² 大阪大,⁽¹³Denali Scientific,⁽¹⁴CNRS・プラズマ物理学研究ユニット,⁽¹⁵UCLA・EPSS,⁽¹⁶UCLA・EPSS,⁽¹⁷ ス ウェーデン王立工科大学,⁽¹⁸ コロラド大学ボルダー校・LASP,⁽¹⁹Southwest Research Institute

Observational evidence of nonlinear growth of whistler-mode waves around quasi-perpendicular bow shocks

#Naritoshi Kitamura¹⁾, Takanobu Amano²⁾, Yoshiharu Omura³⁾, Scott Boardsen^{4,5)}, Masahiro Kitahara⁷⁾, Yoshizumi Miyoshi¹⁾, Satoko Nakamura⁶⁾, Masafumi Shoji¹⁾, Yuto Katoh⁷⁾, Hirotsugu Kojima⁸⁾, Sun-Hee Lee⁹⁾, Daniel J. Gershman⁴⁾, Yoshifumi Saito¹⁰⁾, Masafumi Hirahara¹¹⁾, Shoichiro Yokota¹²⁾, Barbara L. Giles⁴⁾, William R. Paterson⁴⁾, Craig J. Pollock¹³⁾, Olivier Le Contel¹⁴⁾, Christopher Russell¹⁵⁾, Robert J. Strangeway¹⁶⁾, Per-Arne Lindqvist¹⁷⁾, Robert E. Ergun¹⁸⁾, James L. Burch¹⁹⁾

⁽¹ISEE, Nagoya Univ., ⁽²University of Tokyo, ⁽³RISH, Kyoto Univ., ⁽⁴NASA/GSFC, ⁽⁵Goddard Planet. Heliophys. Inst., Univ. of Maryland in Baltimore, ⁽⁶IAR&ISEE, Nagoya University, ⁽⁷Dept. Geophys., Grad. Sch. Sci., Tohoku Univ., ⁽⁸RISH, Kyoto Univ., ⁽⁹Inst. of Astrophys. and Computational Sci., Catholic Univ. of America, ⁽¹⁰ISAS, ⁽¹¹ISEE, Nagoya Univ., ⁽¹²Osaka Univ., ⁽¹³Denali Scientific, ⁽¹⁴LPP, CNRS, ⁽¹⁵EPSS, UCLA, ⁽¹⁶EPSS, UCLA, ⁽¹⁷Royal Institute of Technology, Sweden, ⁽¹⁸LASP, Univ. of Colorado, Boulder, ⁽¹⁹Southwest Research Institute

In this presentation, we show observational results of nongyrotropic electrons close to the cyclotron resonance velocity, which is a key feature of nonlinear growth of whistler-mode waves, using data obtained by the Magnetospheric Multiscale (MMS) spacecraft around quasi-perpendicular bow shocks. Interaction between the electromagnetic field and charged particles is central for the collisionless plasma dynamics in space. Whistler-mode waves are one of the electromagnetic plasma waves, which play important roles in efficient pitch-angle scattering and acceleration of electrons in solar wind, collisionless shock waves as well as planetary magnetospheres. The nonlinear wave-particle interaction theory for coherent large amplitude waves predicts that resonant electrons exhibit nongyrotropy due to phase trapping motion around resonance velocities in the presence of an appropriate inhomogeneity. The nongyrotropic electrons exchange energy and momentum with waves efficiently. While the nonlinear wave growth has been discussed mainly for wave growth in the magnetosphere, it has rarely been discussed for whistler-mode waves in other regions. Here we show examples of nongyrotropic electrons close to the cyclotron resonance velocity during whistler-mode wave events around three quasi-perpendicular bow shock crossings as the observational evidence of nonlinear wave growth. Using measurements by the Fast Plasma Investigation Dual Electron Spectrometer (FPI-DES) and the search-coil magnetometer (SCM), we identified electron flux hole events that induce wave growth; a hole at an appropriate relative phase angle to the whistler-mode wave magnetic field appeared only close to the cyclotron resonance velocity. The magnitudes of the gradient of the magnetic field intensity along the magnetic field line during such time intervals are consistent with an appropriate magnitude to cause phase trapping of resonant electrons. This result provides strong evidence of nonlinear growth of whistler-mode waves around quasi-perpendicular bow shocks. The nonlinear wave growth of whistler-mode waves due to phase trapping, which has been discussed mainly for wave growth in the magnetosphere, probably plays a role in broader applications in space if the appropriate condition is satisfied.