

## Mass-dependent Ion Acceleration in the Plasma Sheet at Storm-time and Non-storm-time Substorms

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It is well known that the flux and energy density of energetic ions increase in the near-Earth plasma sheet during substorm associated dipolarizations. Some studies have focused on the mass dependence of the ion flux increases and found that the flux and energy density of  $O^+$  ions are more enhanced than those of  $H^+$  ions. They have suggested that  $O^+$  ions are accelerated non-adiabatically during dipolarization because the  $O^+$  ion gyro-period is comparable to the time-scale of the magnetic field variations.

Whereas these earlier studies used just a few events for analysis, we analyze 95 events from 10-years of data (from 1995 to 2004) collected by the Energetic Particles and Ion Composition/ Suprathermal Ion Composition Spectrometer (EPIC/STICS) on board the Geotail spacecraft, which measures 9-210 keV/e ions. We compare the rate of  $O^+$  ion energy density before and after a substorm onset to that of  $H^+$  ion energy density. We find that, for 9-36 keV/e ions, the rate of  $O^+$  energy density increase is larger than that of  $H^+$ . However, there is no sharp difference between the two rates for 55-210 keV/e  $O^+$  and  $H^+$  ions. In order to examine whether  $O^+$  ions are accelerated non-adiabatically or not, we compare the gyro-period of  $O^+$  ions ( $T_p$ ) to the time-scale of the magnetic field variation ( $T_m$ ) generated by substorm associated dipolarization. If they were close, the ions should be accelerated non-adiabatically. However, since  $T_m$  is found to be much larger than  $T_p$  in most of the events, we conclude that dipolarization cannot accelerate  $O^+$  ions non-adiabatically and that previous studies observe rare events.

We then separate our events into two groups, storm-time substorms and non-storm-time substorms, and perform the same analysis independently in each group. The main results are as follows: (1) The increasing rate of  $O^+$  with respect to that of  $H^+$  is generally larger during non-storm-time than during the storm-time. (2) The increasing rate of the  $O^+$  energy density in the dusk side of the plasma sheet is larger than that in the dawn side during the non-storm-time. (3) The energy density of low energy  $O^+$  ions increases in the dawn side but decreases in the dusk side during the storm-time.